

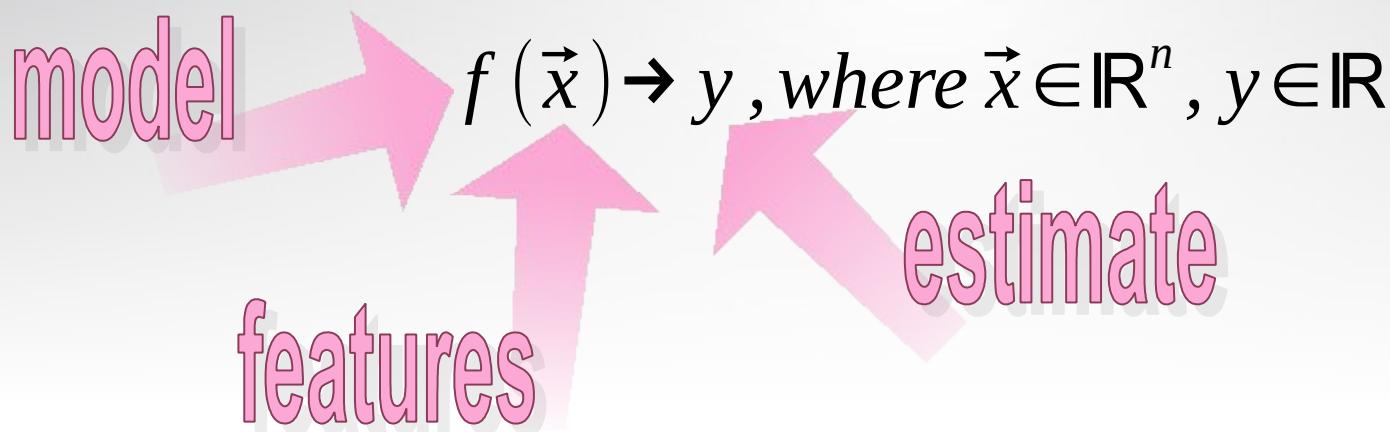
Multimodal Redshift Estimation

proper use of probability density functions

Regression Problems in Astronomy



analysis of
large catalogs demands efficient
solving
regression of problems



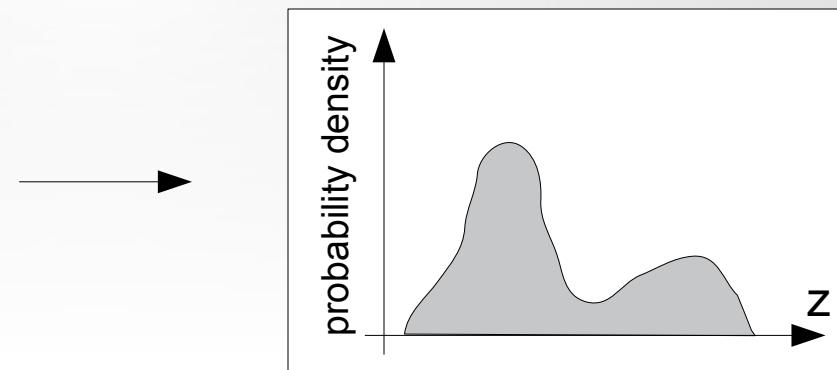
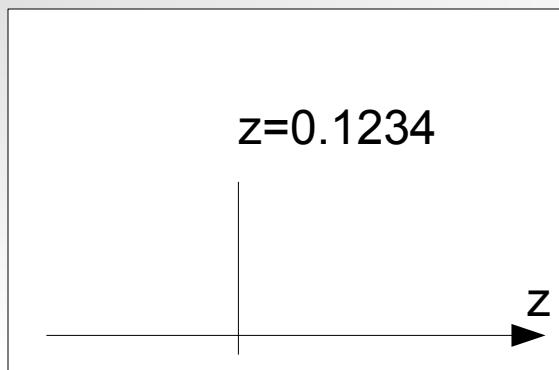
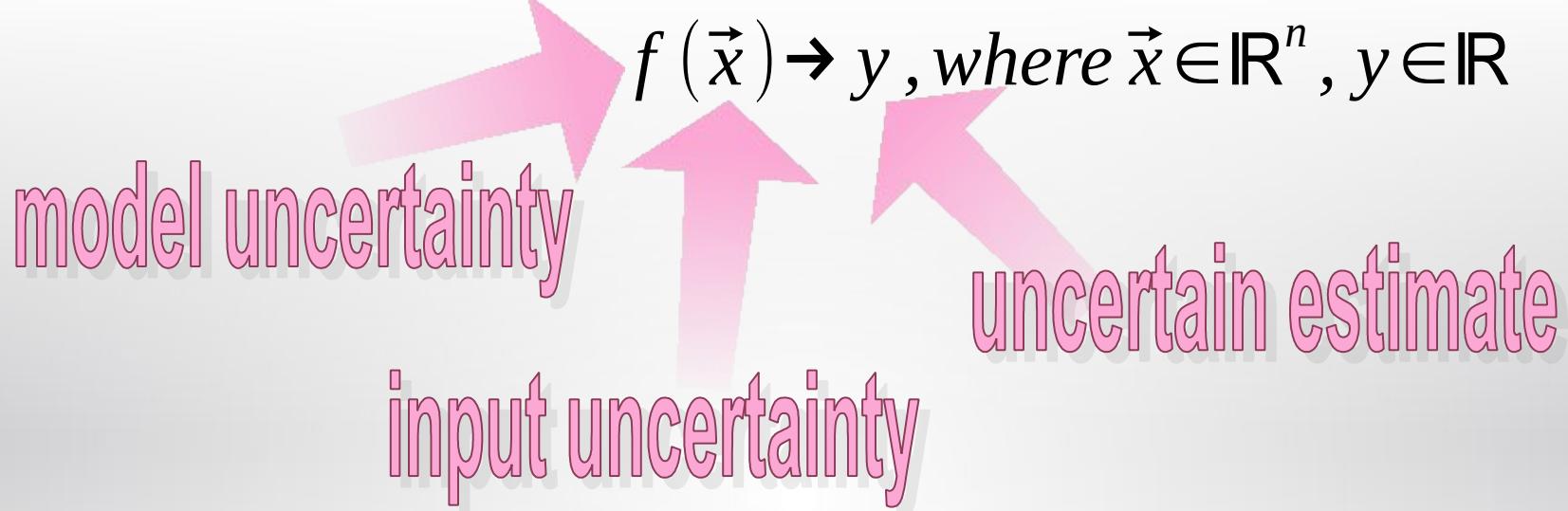
Improvements



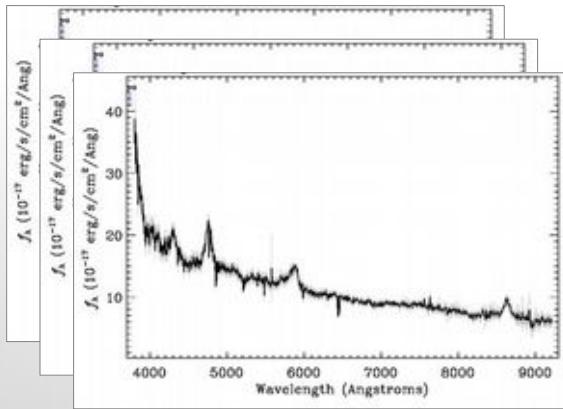
new / improved algorithms
feature / metric selection
domain shift
missing / censored features

how about **uncertainties**?

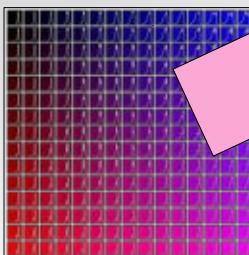
Uncertainties



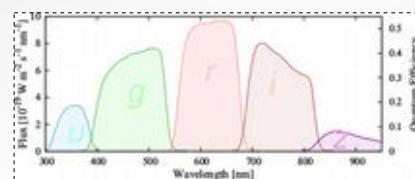
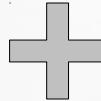
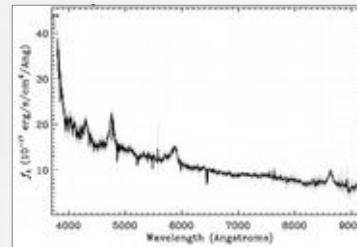
Spectral Model



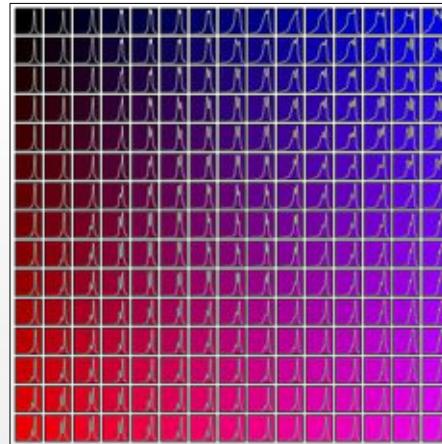
rest frame, $z=0$



shift by z



PPCA



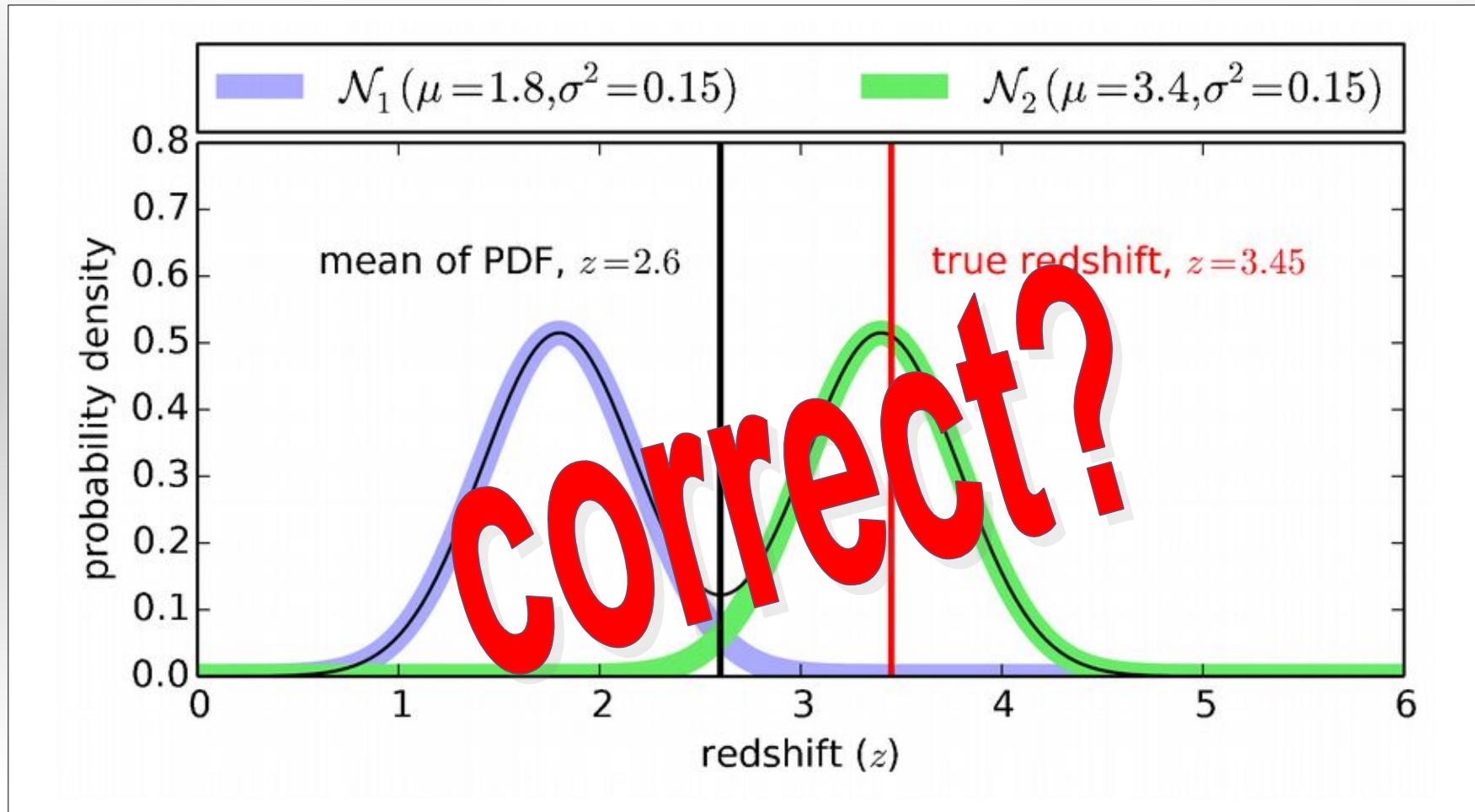
= u,g,r,i,z + noise

compare

observation

Evaluation Tools

simplification / just use the mean



Evaluation Tools

stacking PDFs



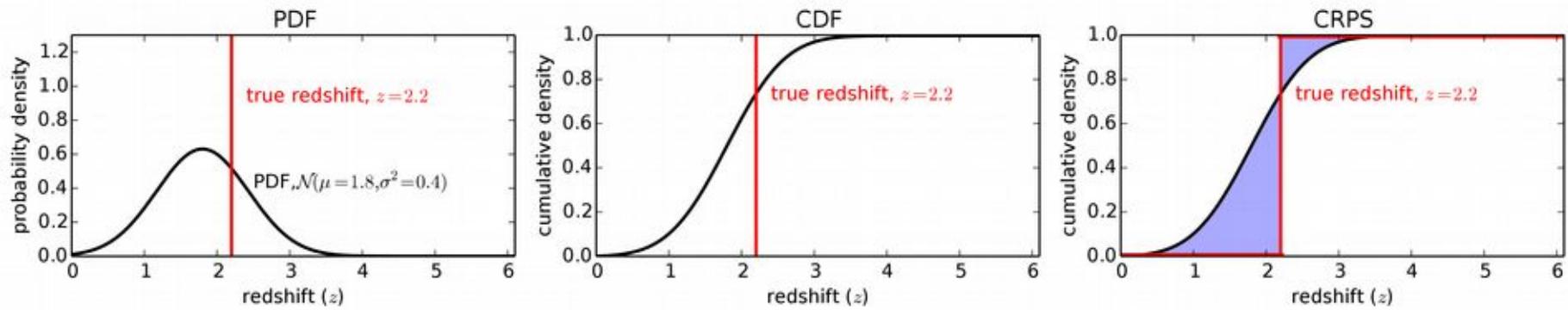
Proper Evaluation Tools / CRPS



continuous rank probability score

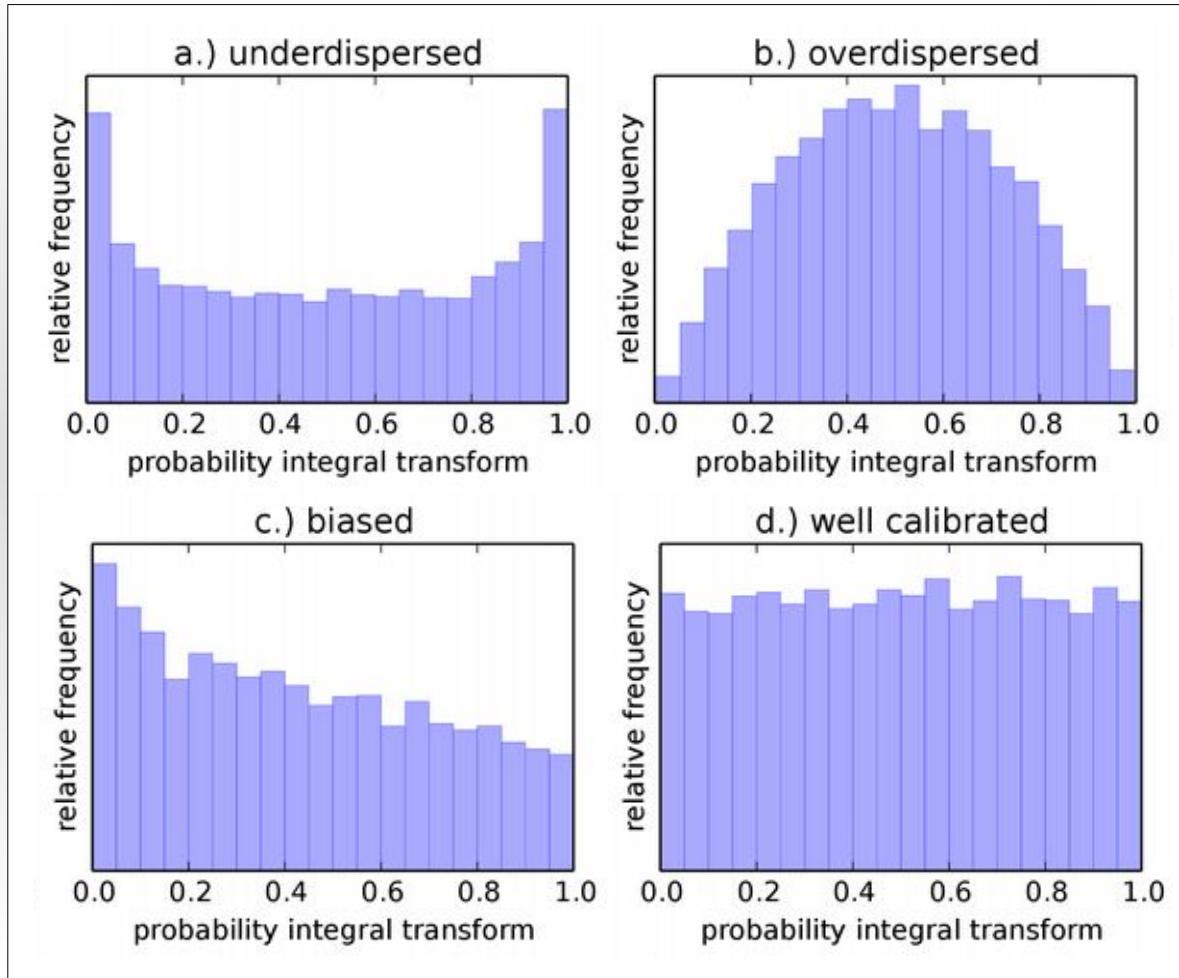
$$CRPS = \frac{1}{N} \sum_{t=1}^N crps(CDF_t, z_t),$$

$$\text{with } crps(CDF_t, z_t) = \int_{-\infty}^{+\infty} [CDF_t(z) - CDF_{z_t}(z)]^2 dz$$

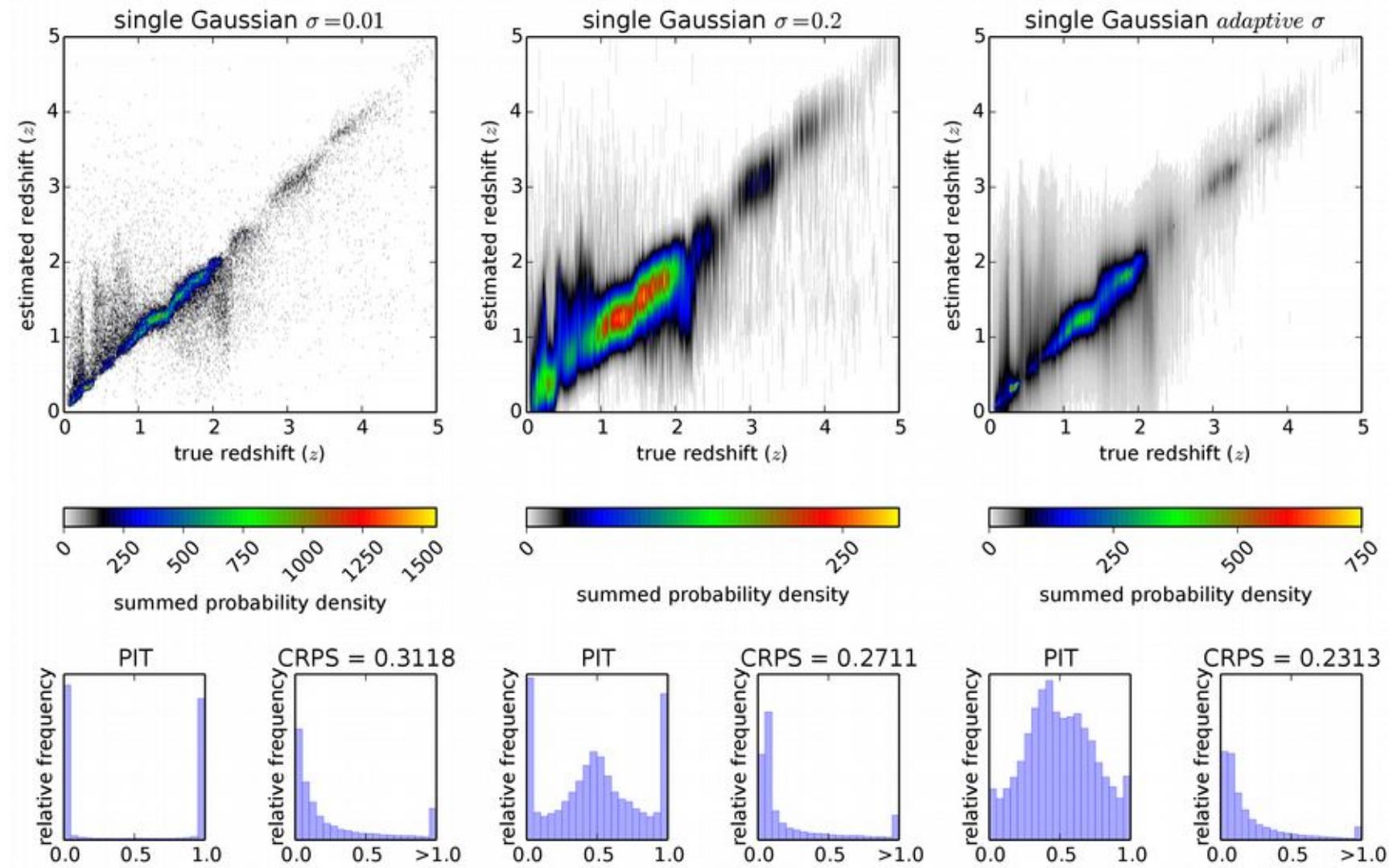


Proper Evaluation Tools / PIT

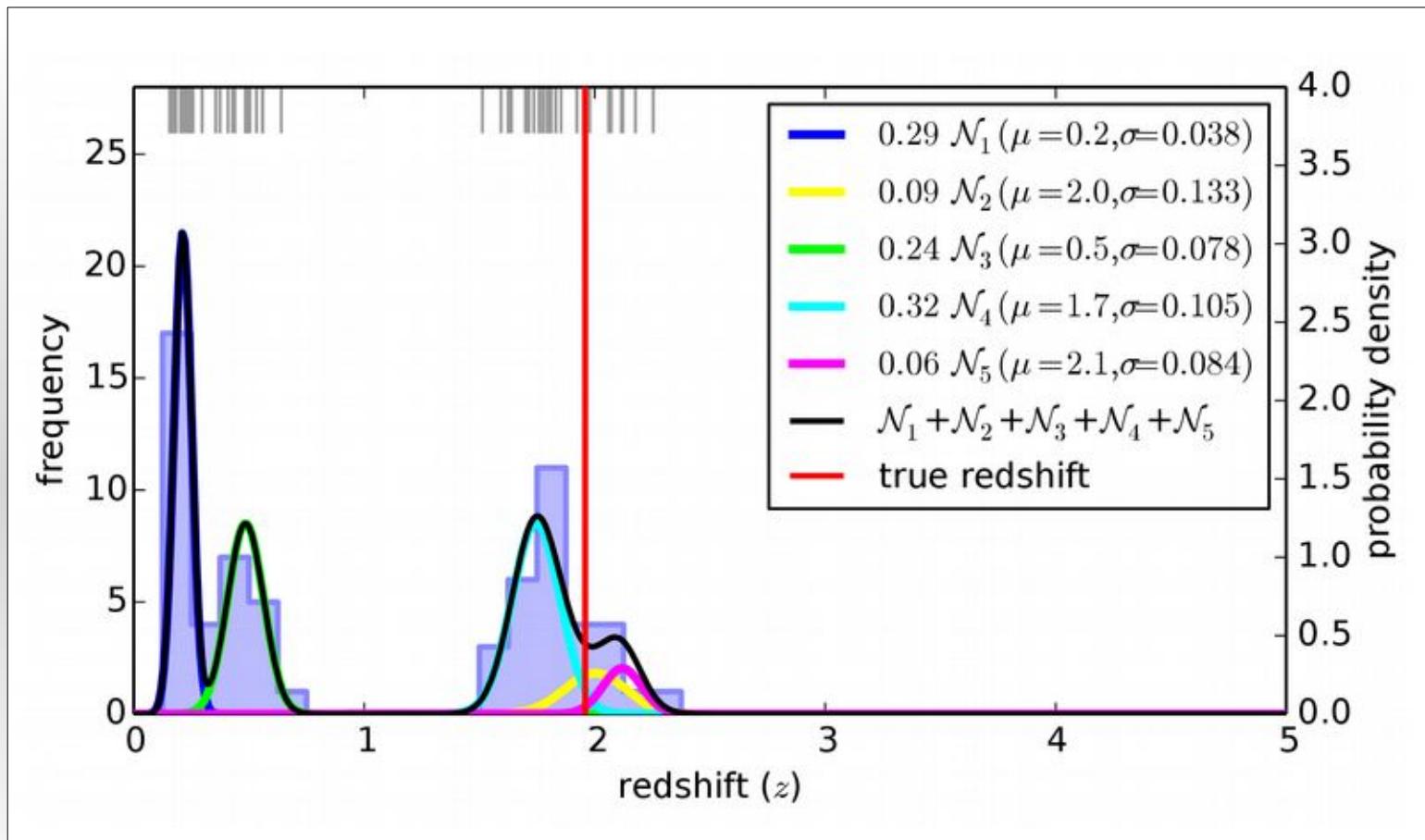
probability integral transform



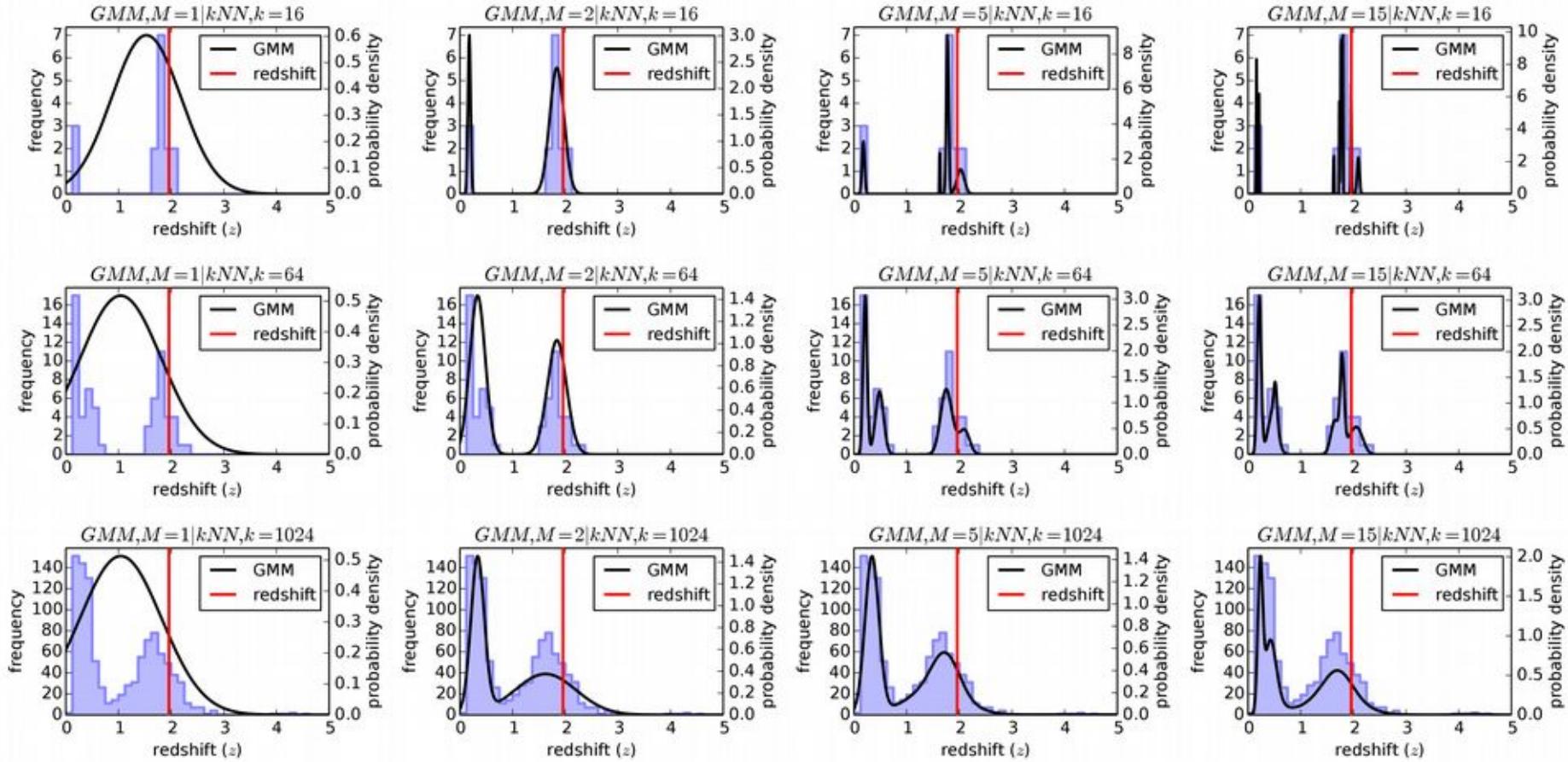
Results



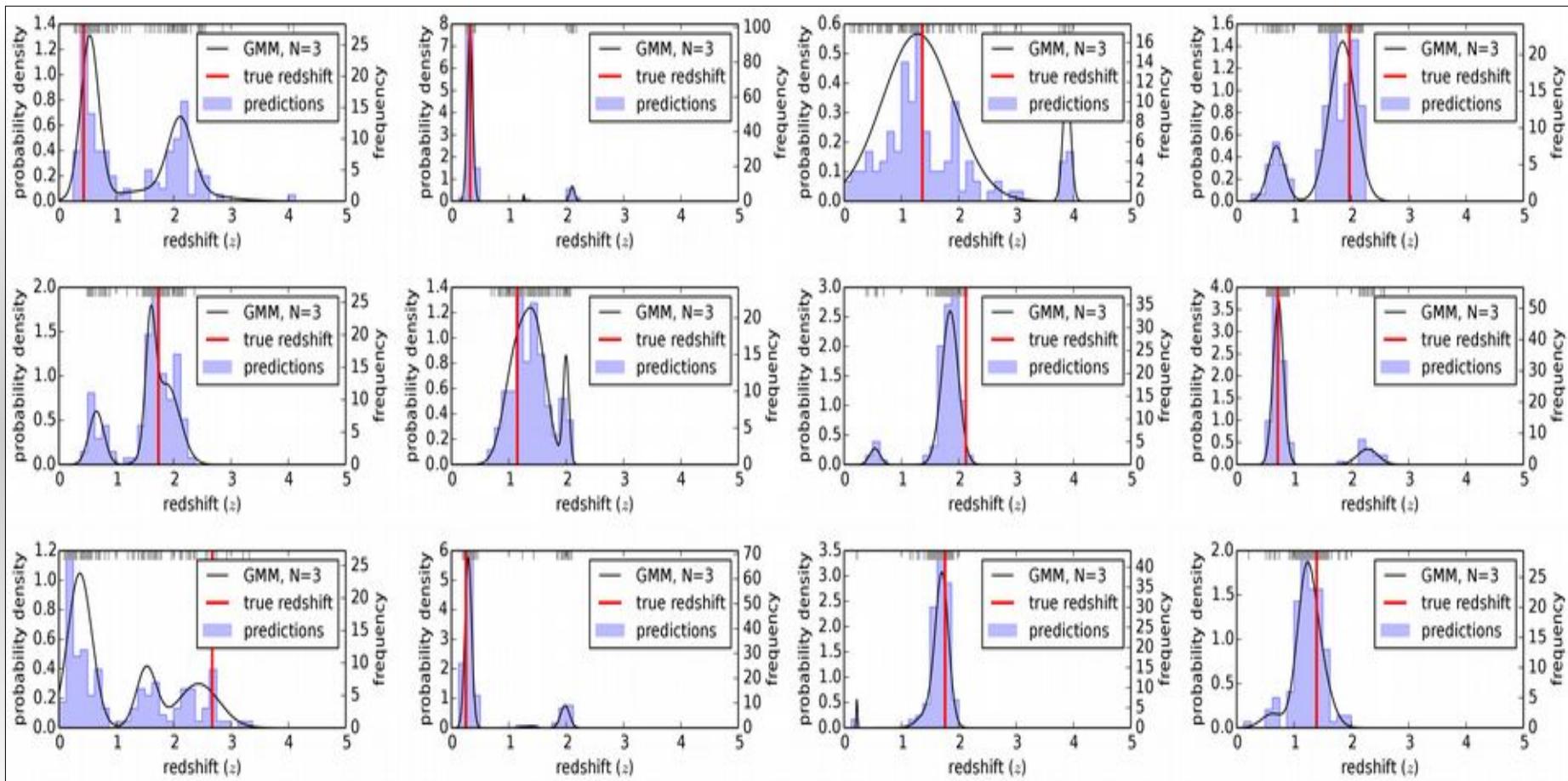
Multi-Modalities



Multi-Modalities



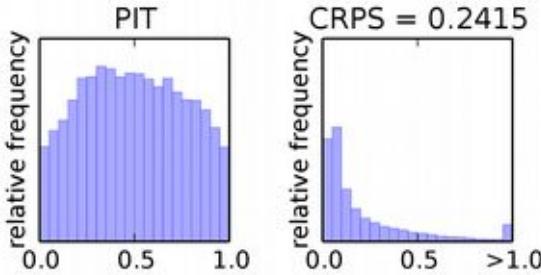
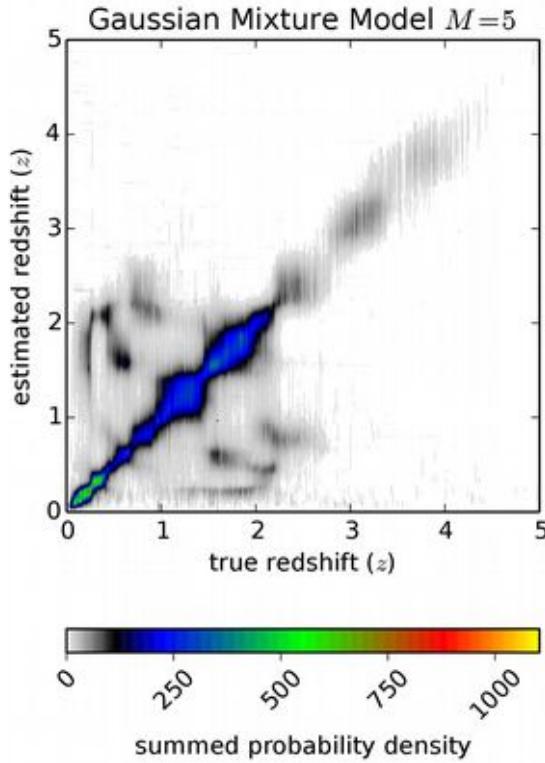
Multi-Modalities



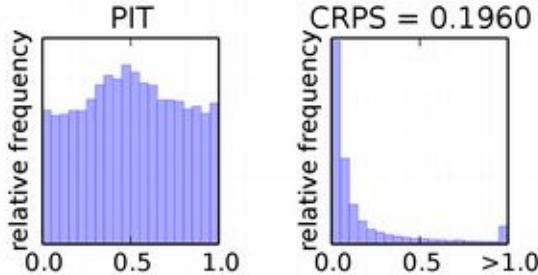
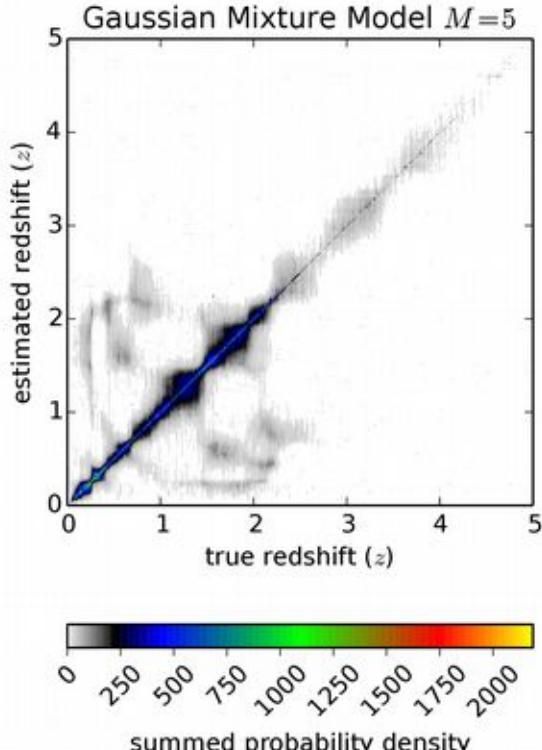


Results

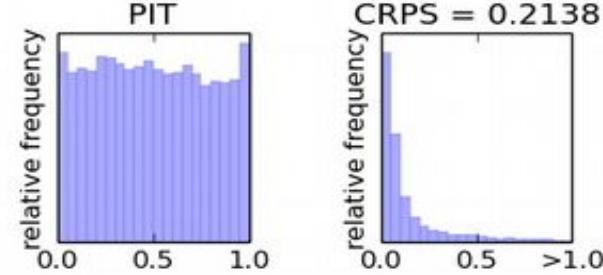
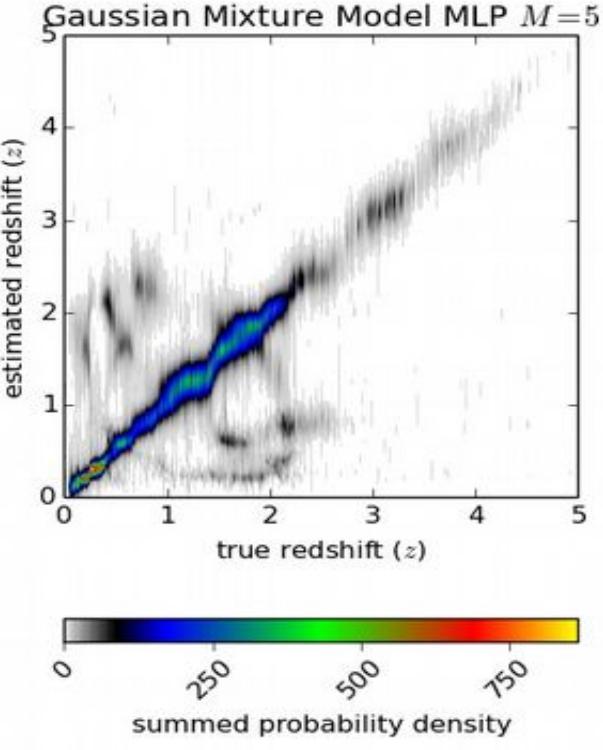
Nearest Neighbors



Random Forest

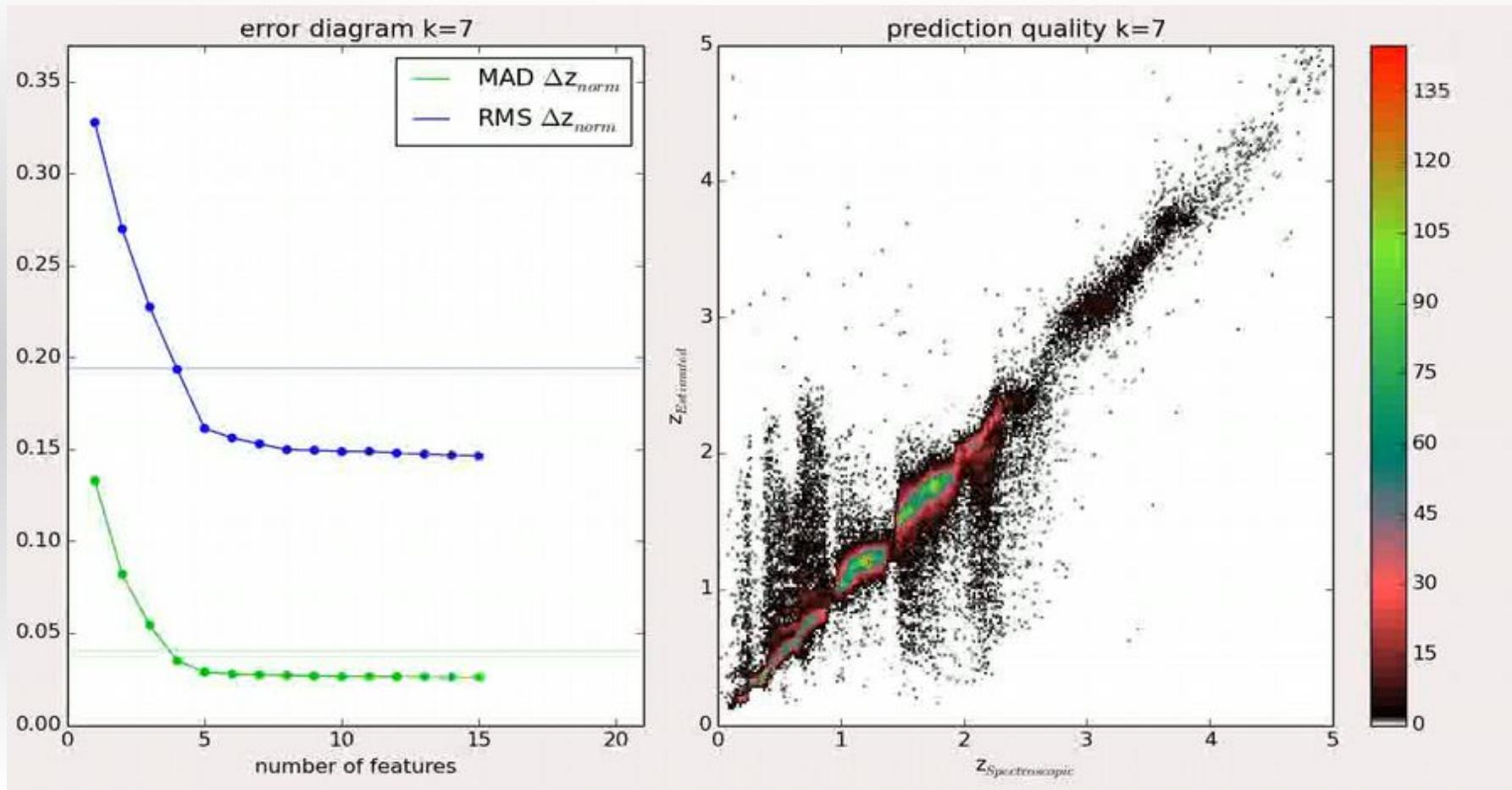


Mixture Density Network



Forward Selection

apply greedy forward selection



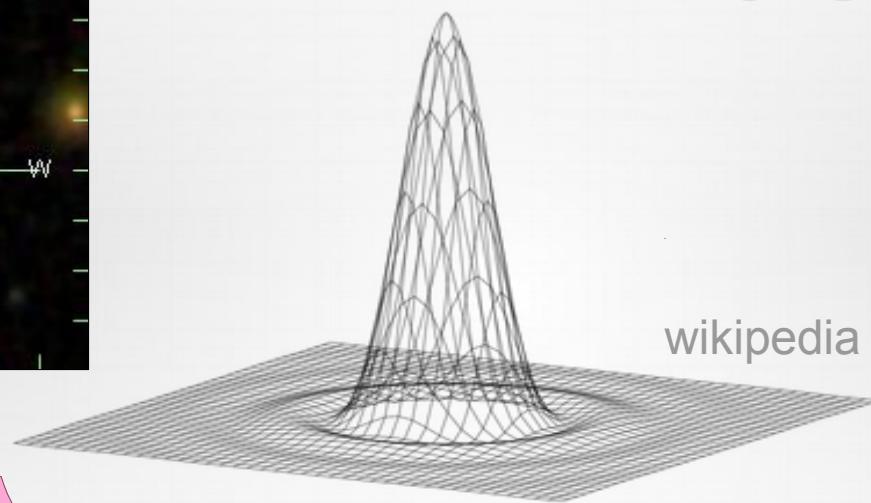
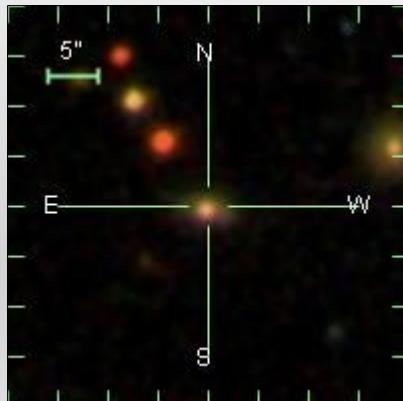
Forward Selection

resulting features:

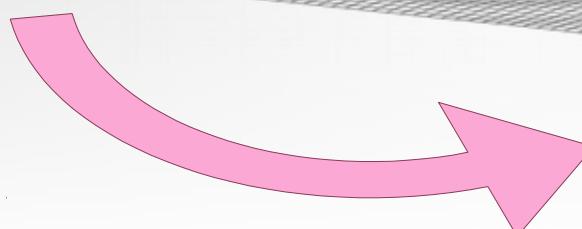
$$\frac{u_{psf} - g_{petrosian}}{dered(z_{psf}) - dered(i_{petrosian})}$$
$$\frac{dered(g_{psf}) - dered(r_{model})}{dered(r_{psf}) - dered(z_{model})}$$
$$\sqrt{\sigma_{g_{model}}^2 + \sigma_{r_{model}}^2}$$
$$dered(r_{model}) - dered(i_{model})$$
$$\frac{i_{psf} - i_{petrosian}}{dered(z_{psf}) - dered(r_{petrosian})}$$
$$\frac{g_{model} - g_{petrosian}}{\sqrt{\sigma_{g_{petrosian}}^2 + \sigma_{r_{petrosian}}^2}}$$

Lessons Learned

features stored in catalogs
are not the best for machine learning

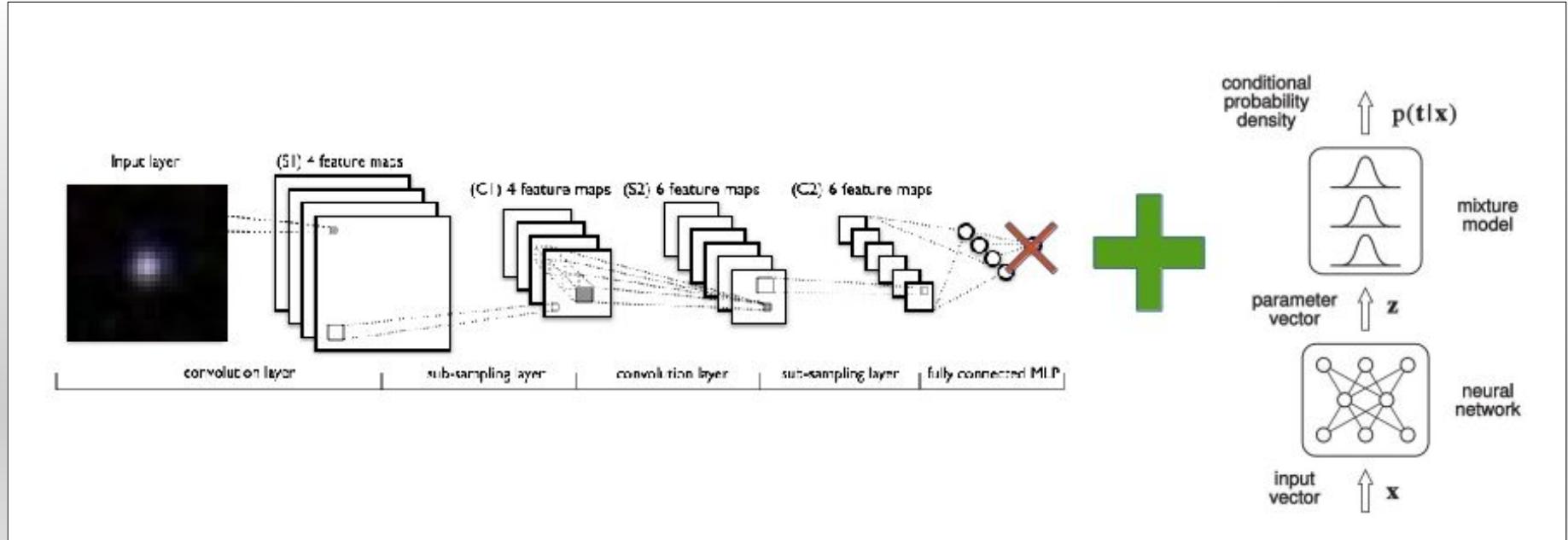


wikipedia

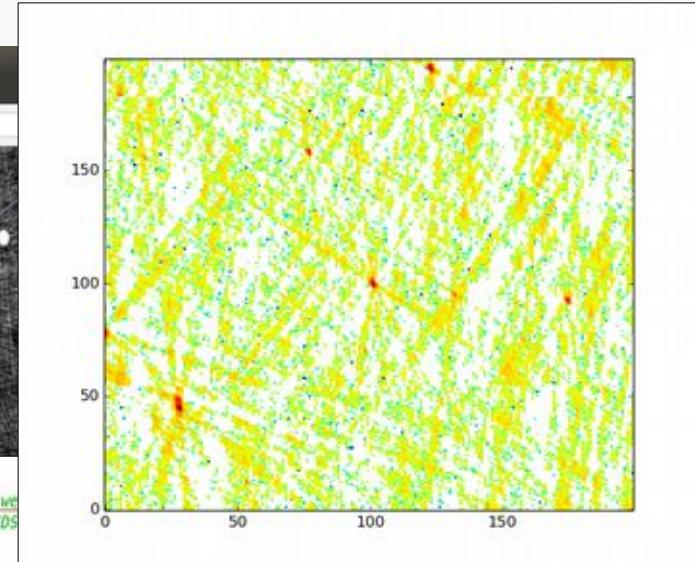
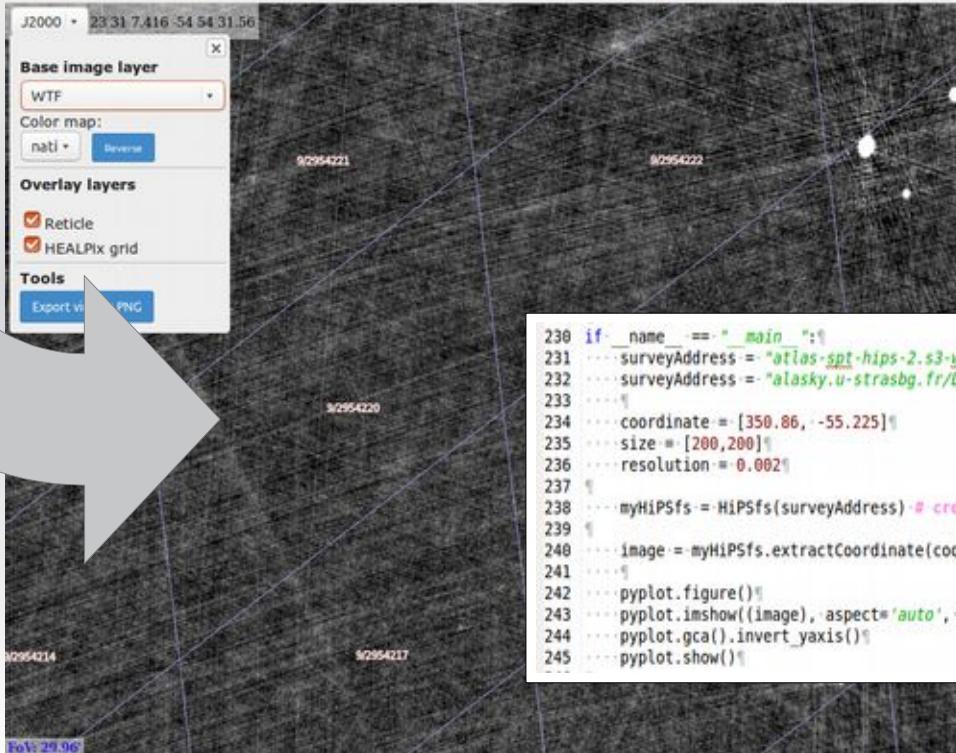
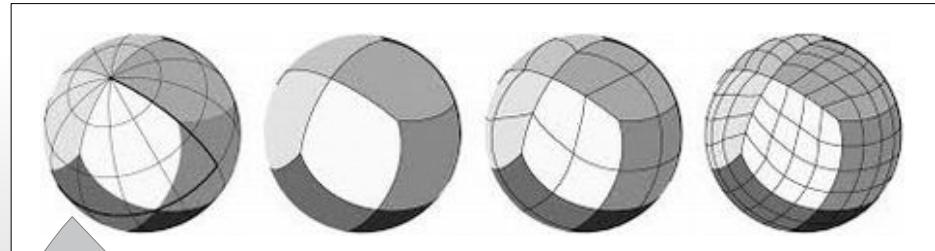


magnitudes
in (u, g, r, i, z)

DCN meet MDN



Healpix / HiPS / IVOA

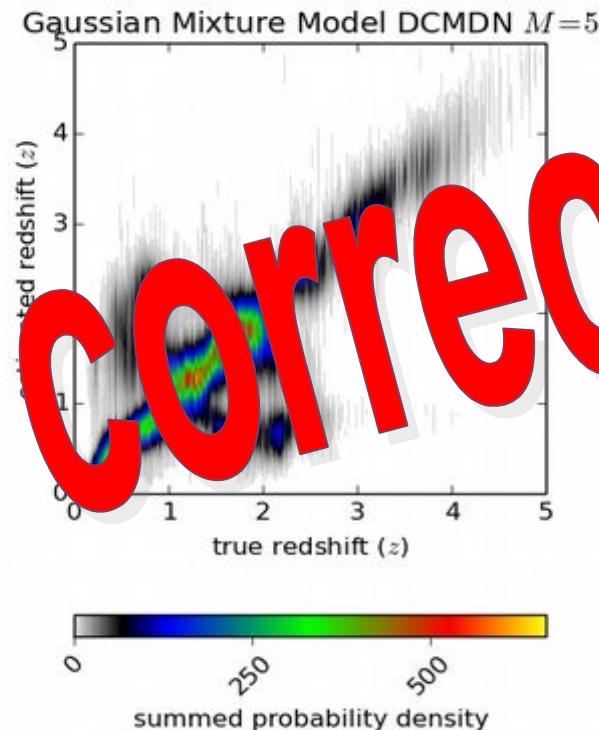


```
230 if __name__ == "__main__":
231     surveyAddress = "atlas-spt-hips-2.s3-west-1.amazonaws.com"
232     surveyAddress = "alasky.u-strasbg.fr/DS"
233     ...
234     coordinate = [350.86, -55.225]
235     size = [200,200]
236     resolution = 0.002
237
238     myHiPSfs = HiPSfs(surveyAddress) # create access
239
240     image = myHiPSfs.extractCoordinate(coordinate, size, resolution, nested=True) # extract data array
241
242     pyplot.figure()
243     pyplot.imshow(image, aspect='auto', interpolation="nearest")
244     pyplot.gca().invert_yaxis()
245     pyplot.show()
```

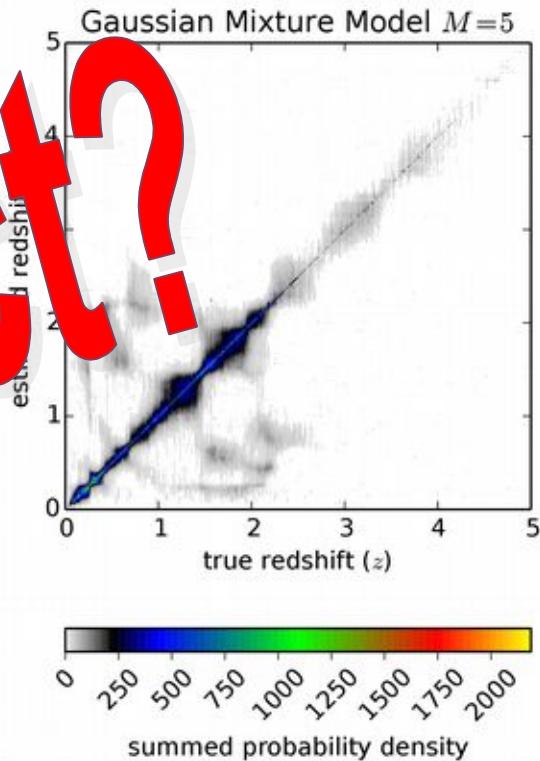


Results

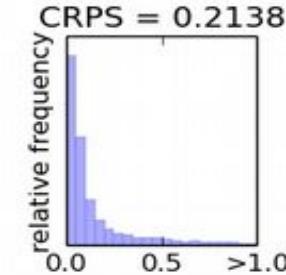
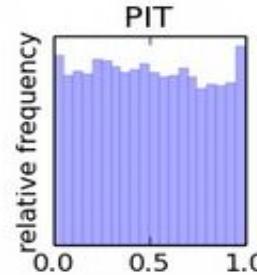
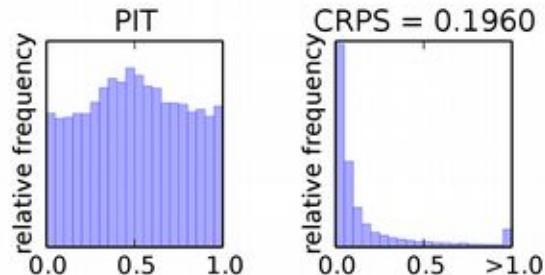
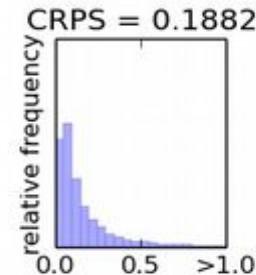
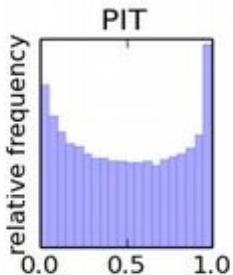
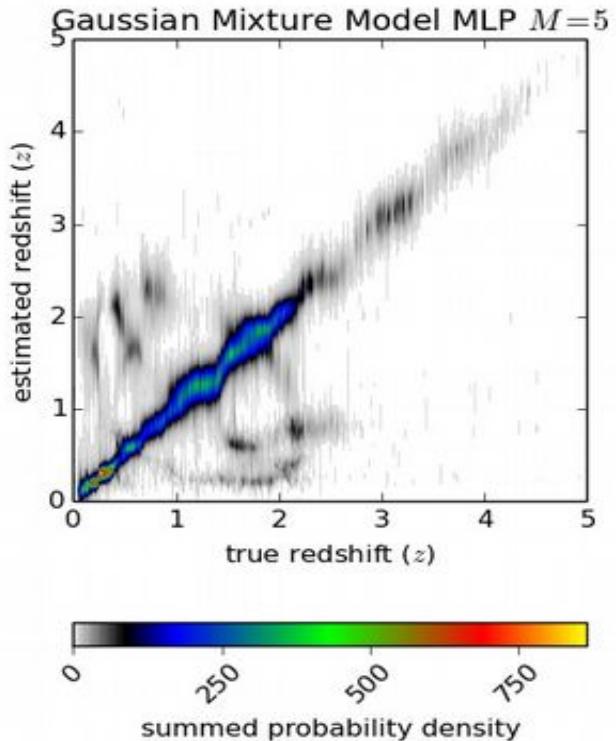
DCN+MDN



Random Forest



Mixture Density Network



Conclusion



uncertainty is very
important to further
improve
the science quality

we must start to use
proper
tools to
evaluate the performance